

The Issac Newton Telescope Monitoring Survey of Local Group Dwarf Galaxies VIII, The Star Formation History and Dust Production Rate of Andromeda Dwarf Satellites

Hedieh Abdollahi¹, Atefeh Javadi¹, Jacco Th. van Loon², Iain McDonald^{3, 4}, Habib G. Khosroshahi¹, Bernard Foing^{5, 6}, Fatemeh Fazel⁵, Hamidreza Mahani¹, and Mahdi Abdollahi¹

¹School of Astronomy, Institute for Research in Fundamental Sciences (IPM), Tehran, Iran

²Lennard-Jones Laboratories, Keele University, UK

³Jodrell Bank Centre for Astrophysics, Alan Turing Building, University of Manchester, UK

⁴Department of Physical Sciences, The Open University, Walton Hall, Milton Keynes, UK

⁵Leiden Observatory, Leiden University, PO Box 9513, 2300 RA Leiden, The Netherlands

⁶ESTEC, ESA, Keplerlaan 1, 2201 AZ Noordwijk, The Netherlands

March 12, 2024

An extensive analysis of observed spheroidal dwarf satellites of the Andromeda galaxy was conducted to provide a comprehensive catalog of long-period variable (LPV) stars. The study is part of an optical monitoring survey to observe 55 dwarf galaxies and four globular clusters within the Local Group. The dwarfs were observed using the wide field camera (WFC) of the 2.5 meter Isaac Newton Telescope (INT). These observations were conducted through near-infrared (*i* Sloan) and optical (*V* Harris) bands at nine epochs. Our focus is on detecting asymptotic giant branch (AGB) stars when their magnitude amplitude becomes higher than 0.2 mag to track dwarf evolution. In dwarf galaxies, LPV stars with ages ranging from 30 Myr to 10 Gyr provide powerful tools for reconstructing the star formation history (SFH). LPV stars are suitable for study due to the relation between their birth mass and luminosity at the end of their evolution. Star formation scenarios could also be studied by comparing the age gradient of variable stars between two half-light radii. This study also provides estimates of satellite characteristics such as the half-light radius, the

RGB-tip, the distance modulus, the quenching time, and the total stellar mass. Our work is extended by estimating the mass-loss rates of detectable LPVs to produce dust maps for each dwarf. Mass-loss of LPVs contributes to galaxy chemical enrichment and star formation. A combination of ground-based observations and space-based telescopes was used to simulate the spectral energy distribution (SED) for LPVs based on observations in the optical to mid-infrared bands. It is also possible to estimate the re-ignition time for SFH by calculating the total mass return of LPVs. Additionally, the distance modulus has been updated using RGB-tip through the tip detection of the Sobel filter in this study. The distance modulus of Andromeda satellites ranges from 23.57 ± 0.08 to 25.62 ± 0.17 mag.